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# Deep learning methods for gamma event selection in TAIGA-IACT image analysis in stereo-mode

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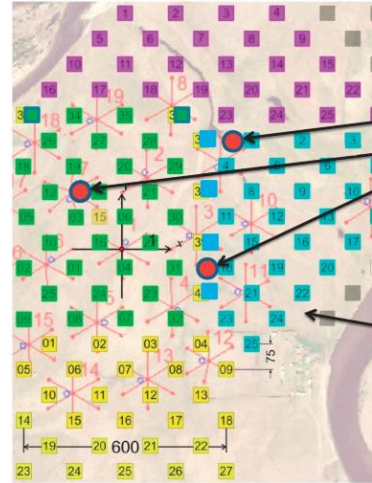
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# TAIGA-IACT

- Imaging Air Cherenkov Telescopes (IACT) are located in Tunka Valley, Republic of Buryatia, at the astrophysical gamma observatory TAIGA of the API ISU.
- Mirror diameter: 4.3m
- FoV:  $9.6^\circ$
- Camera: 600 PMT XP1911, grouped into 22 clusters.
- Four TAIGA-IACTs are in operation mode.

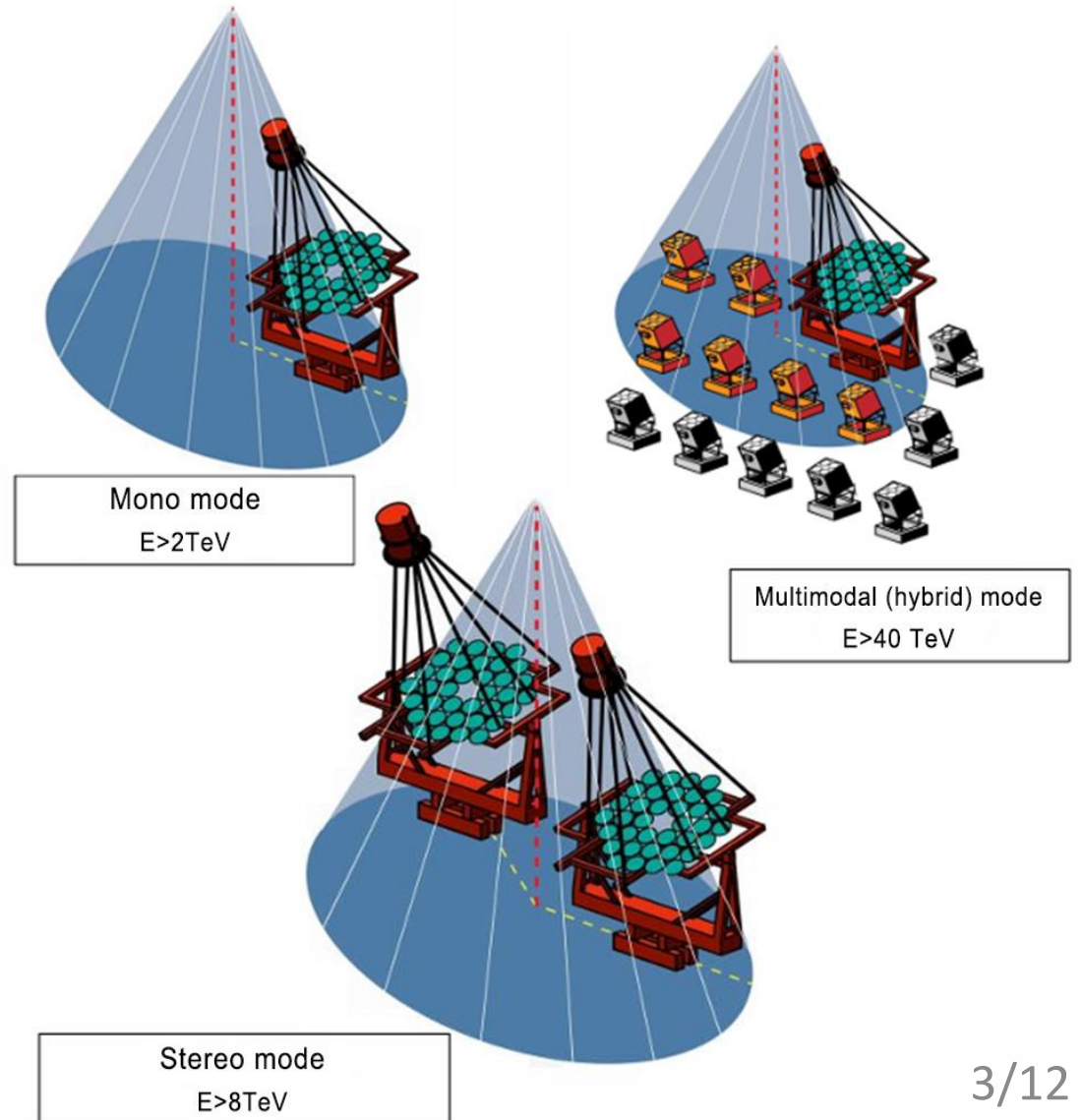
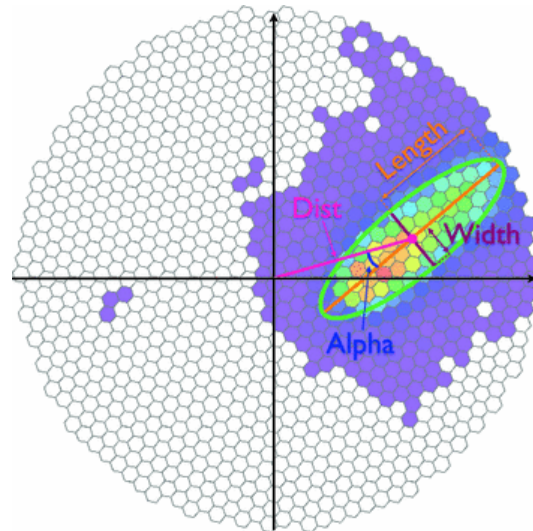


$51^\circ 48' 35'' \text{ N}$   
 $103^\circ 04' 02'' \text{ E}$   
675 m a.s.l.



# Stereo mode

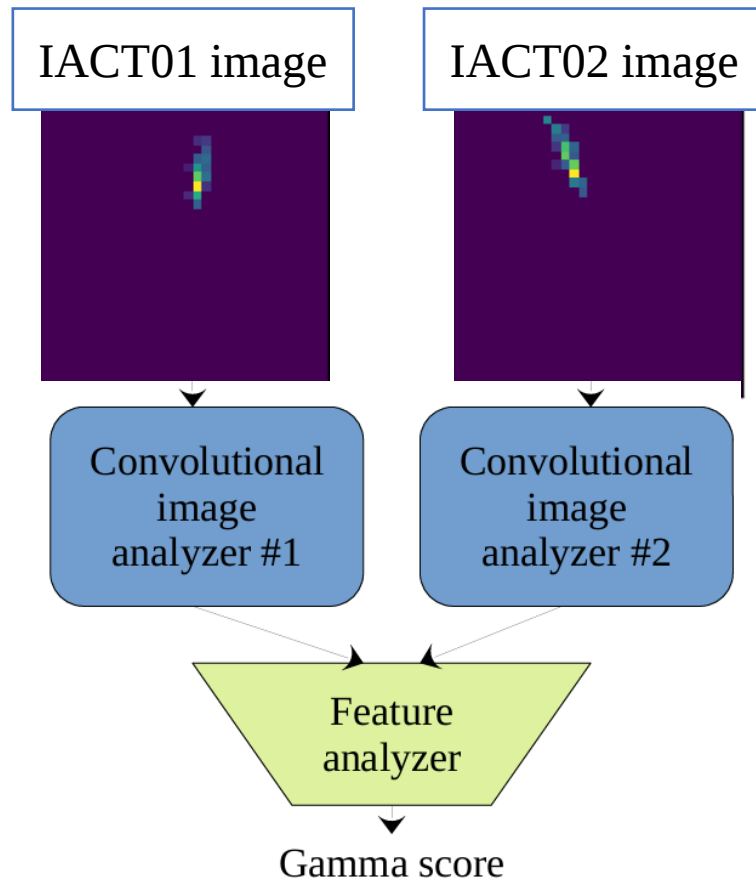
- Every telescope detects the Cherenkov light from EAS.
- Since telescopes are located near each other (250-350 m), Cherenkov light cone can trigger multiple telescopes.
- Classical approach is based on the analysis of normalized Hillas parameters extracted from the images.



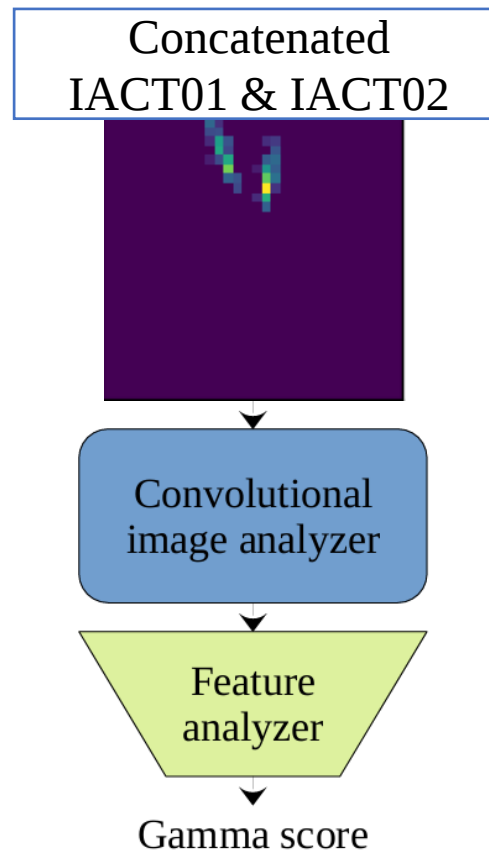
# Proposed deep learning approaches for IACT stereo-mode

- The aim of the research: develop neural network methods for gamma selection with TAIGA-IACTs in stereoscopic observations

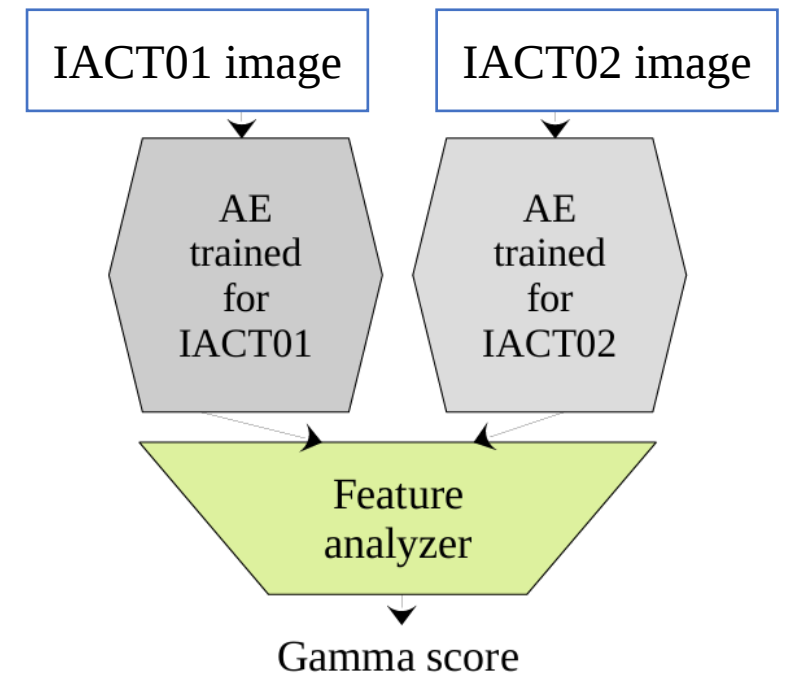
## 1. Medium fusion:



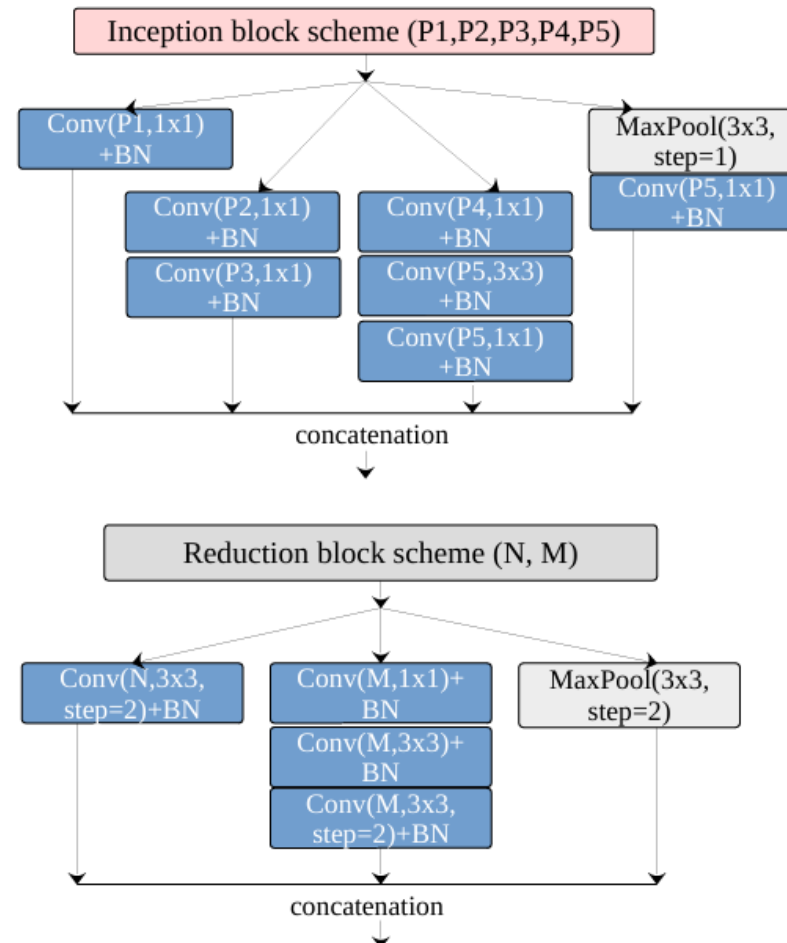
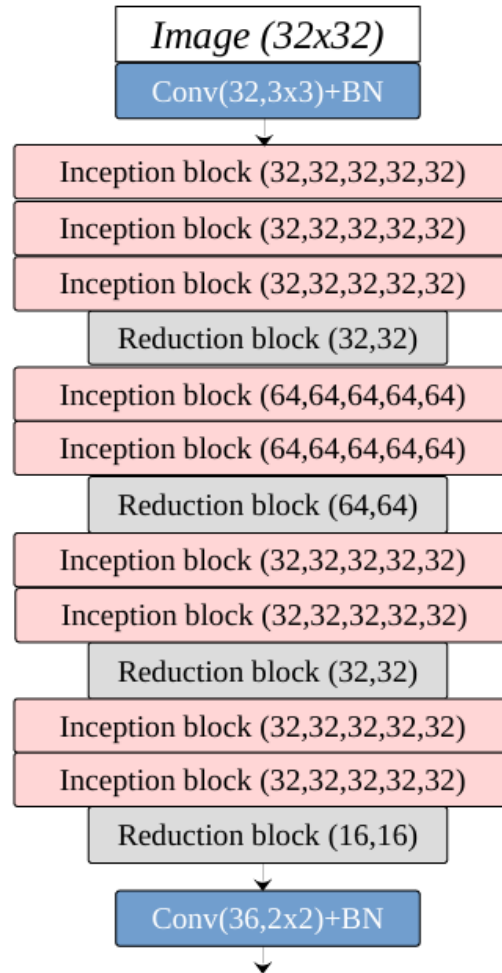
## 2. Early fusion:



## 3. Essential feature analysis

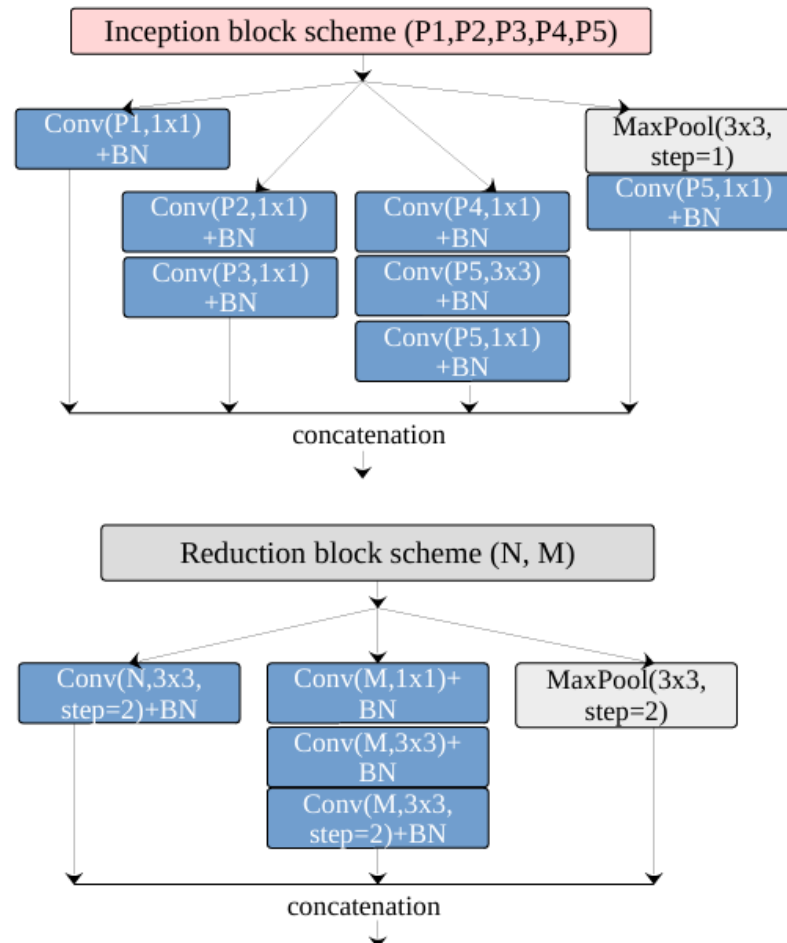
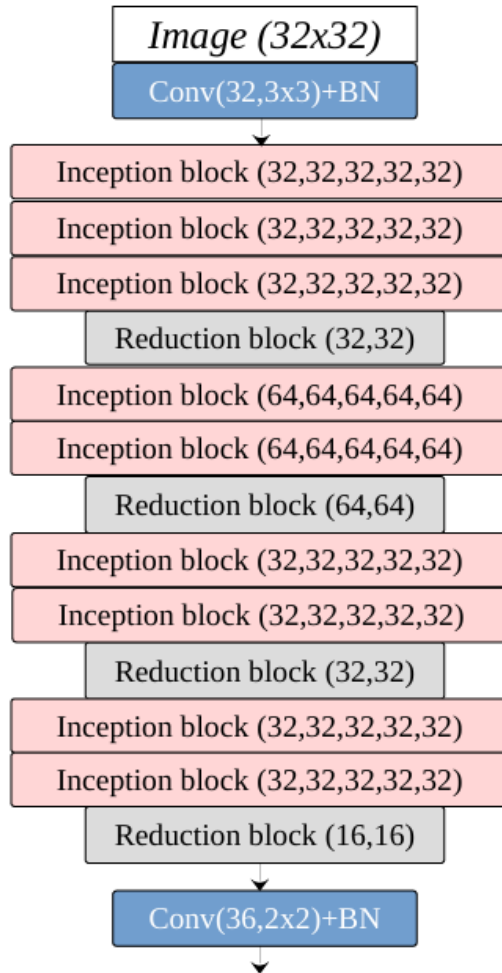


# Proposed deep learning approaches for IACT stereo-mode



- Simplified Inceptionv4 CNN
- Total trained weights: ~1.1 million
- In case of medium and early image fusion, 1 dense layer is feature analyzer
- For AE latent space analysis feature analyzer structure:
- Input(36) -> dense(2048) -> reshape(8,8,32)-> conv2d(2x2,128, s=2) -> conv2d(2x2,256, s=2) -> conv2d(2x2,512, s=2) -> dense(512) -> dense(1)

# Proposed deep learning approaches for IACT stereo-mode

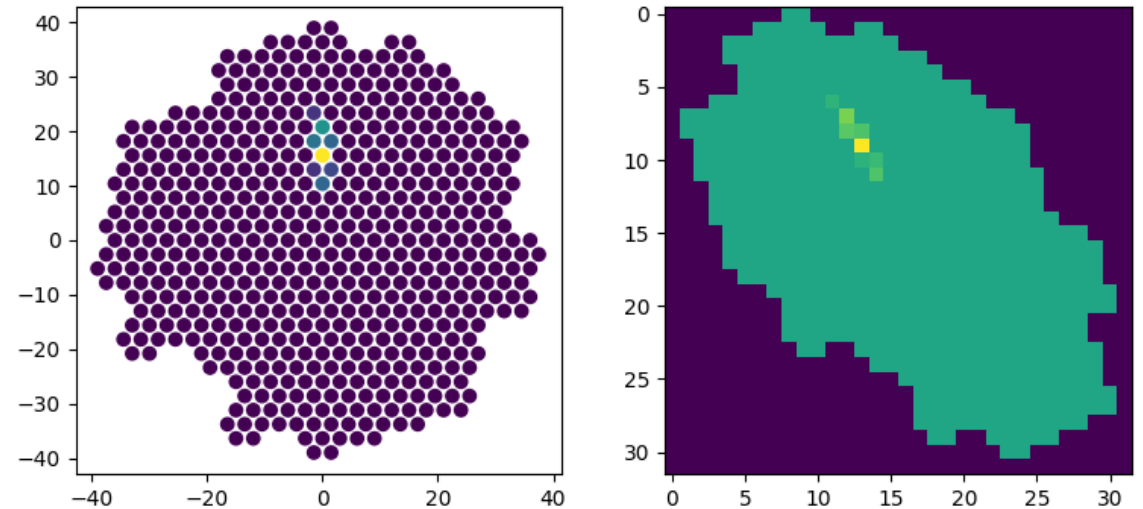


- Classification loss function: binary crossentropy, *focal loss*.
- Image reconstruction loss function: MSE.
- Activation function: silu
- Epoches (classifier): ~10
- Epoches (AE): ~100
- Learning rate: 0.0001, with step decay

# Training and Validation

- MC data was generated with CORSIKA. Simulation includes relative PMT sensitivity, PMT's Winston cone, background noise, etc.
  - Gamma and proton events spectrum slope: -1
  - Energy range: 40-400 TeV.
- Gamma training events: 29 900
- Proton training events: 157 000
  - AE was trained on proton events.
  - Augmentation for stereo-mode CNN: class weights, input channel switch.

- Image preprocessing:
  - Image cleaning;
  - Image transformation to square form;
  - Logarithmic amplitude scaling.

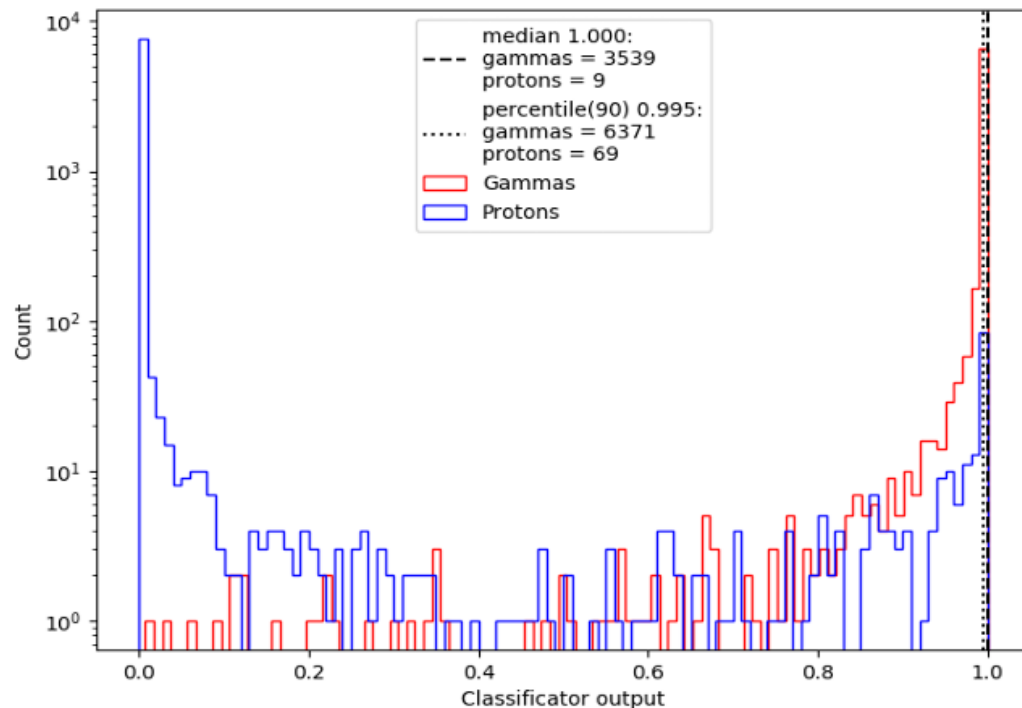


Example of image transformation

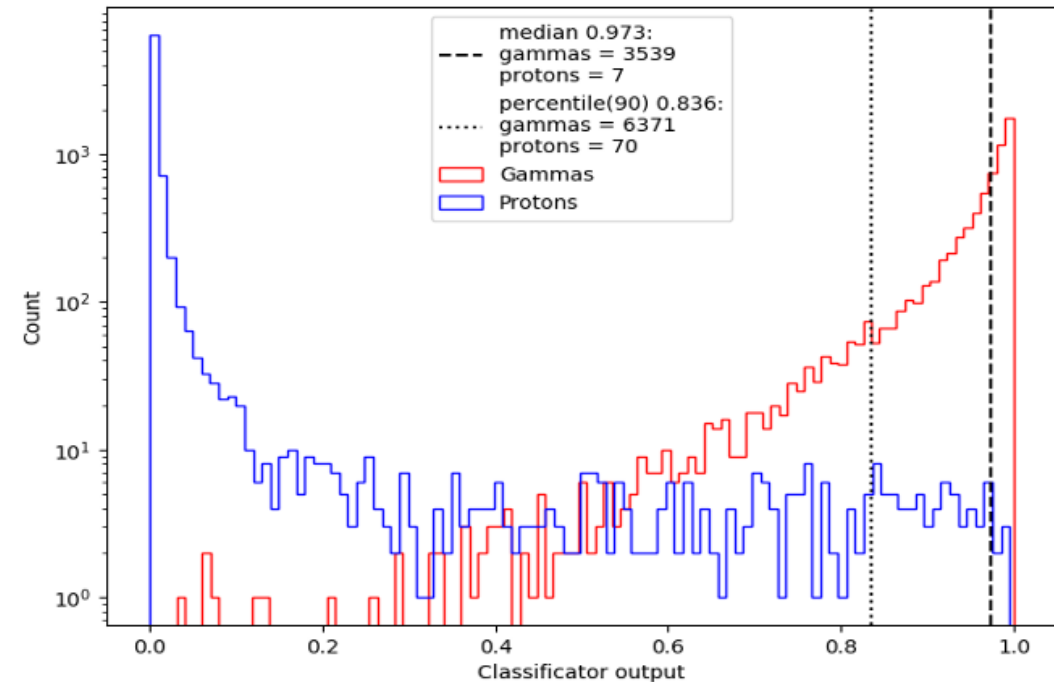
# Training and Validation

- Validation gamma events: 7 100
- Validation protons events: 7 600
- Class threshold: 50% gamma remaining

Binary crossentropy (0.99994)

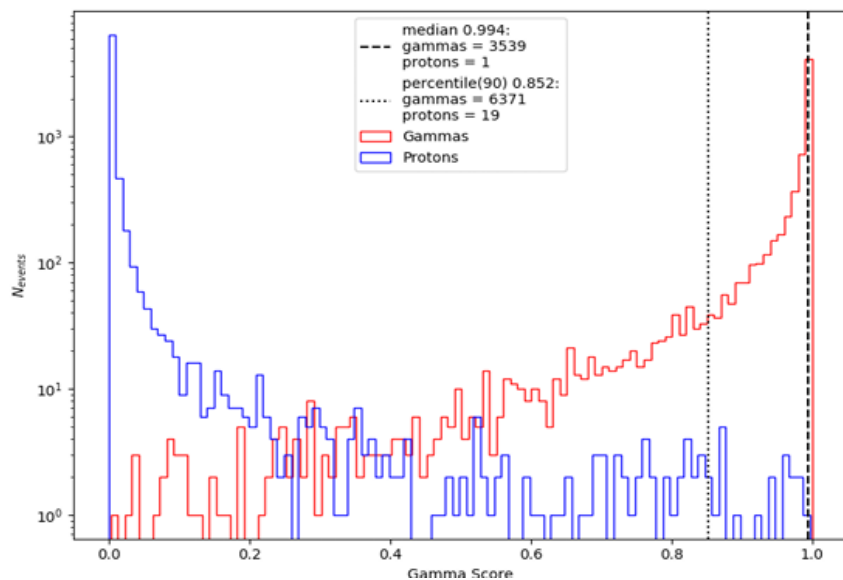


Focal loss (0.973)

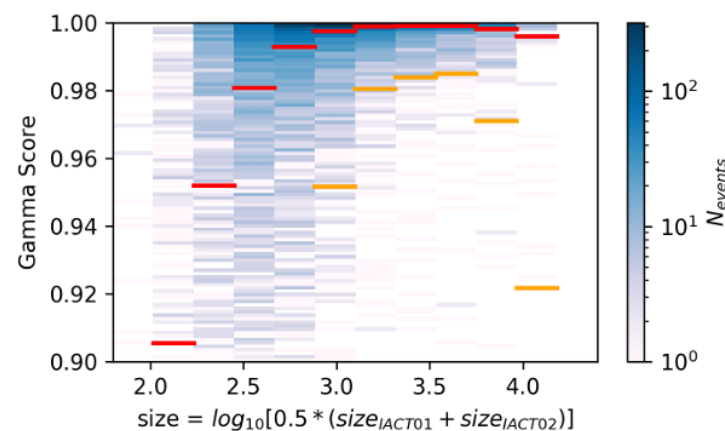


# Training and Validation

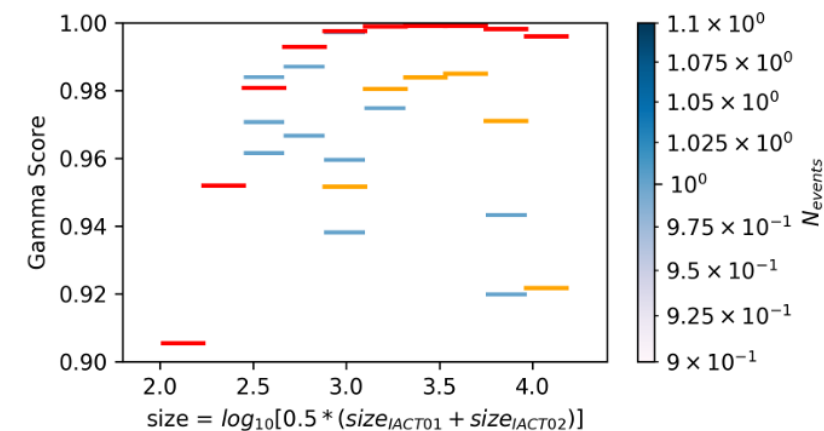
Medium fusion



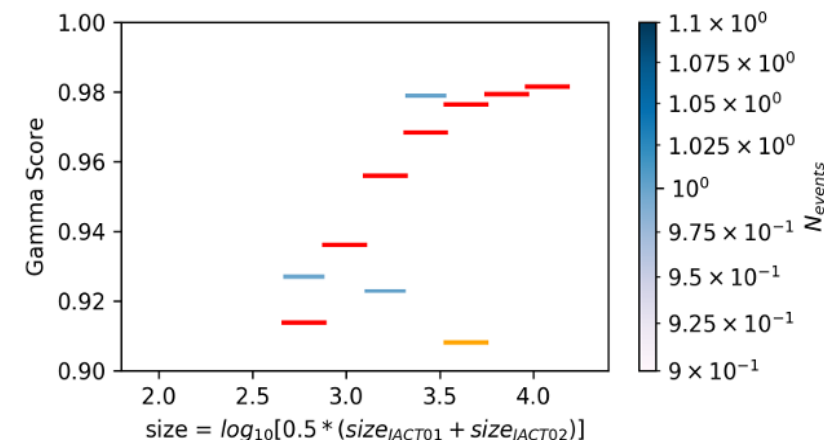
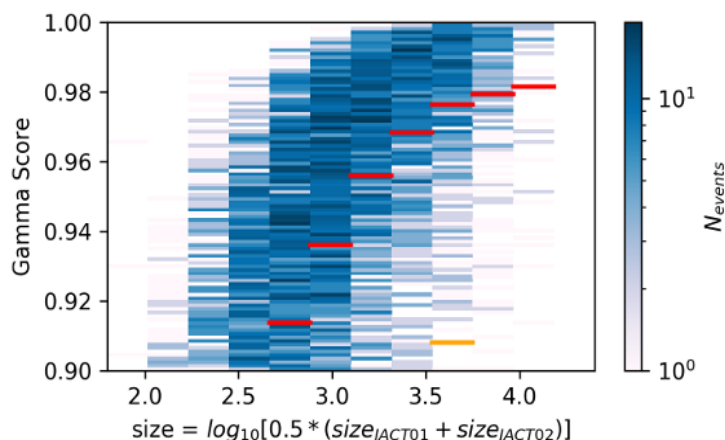
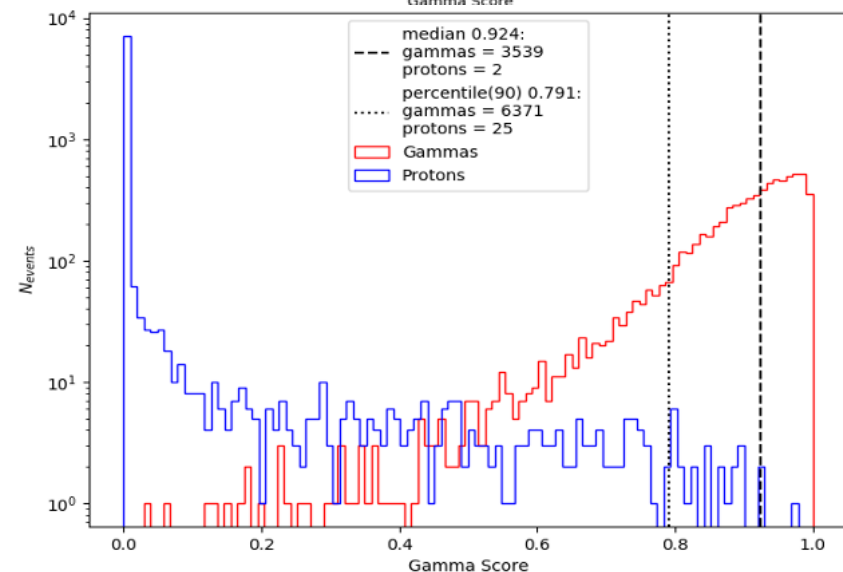
Gamma



Protons



Latent space analysis



On pictures above: red – 50% gamma threshold, yellow – 90% gamma

## Selection in case of class imbalance

- For the most gamma sources, significant gamma/hadron imbalance (1/10000) are observed in TeV energy range.
- Thus, test imbalance was around 1/100 to investigate the CNN capabilities. Also, test results for mono-mode are presented. Class thresholds obtained from CNN validations.

|                                  | DL methods                    | Total gamma | Total hadrons | selected gamma | selected protons | Gamma remained | Hadron suppression |
|----------------------------------|-------------------------------|-------------|---------------|----------------|------------------|----------------|--------------------|
| Size-independent class threshold | CNN (mono-mode)               | 320         | 35000         | 147            | 41               | 45.9%          | 854                |
|                                  | CNN (medium fusion)           | 400         | 39600         | 188            | 9                | 47.0%          | 4400               |
|                                  | CNN (early fusion)            | 400         | 39600         | 206            | 3                | 51.5%          | 13200              |
|                                  | AE essential feature analysis | 400         | 39600         | 199            | 16               | 49.8%          | 2475               |
| Size-dependent class threshold   | CNN (mono-mode)               | 320         | 35000         | 151            | 119              | 47.2%          | 294                |
|                                  | CNN (medium fusion)           | 400         | 39600         | 188            | 26               | 47.0%          | 1523               |
|                                  | CNN (early fusion)            | 400         | 39600         | 203            | 18               | 50.8%          | 2200               |
|                                  | AE essential feature analysis | 400         | 39600         | 217            | 40               | 54.3%          | 990                |

## Conclusion

- Multiple deep learning methods was proposed and studied in the gamma/hadron classification task: medium fusion (separate IACTs images analysis in CNN before concatenation), early fusion (concatenated IACTs images are fed to CNN) and autoencoder-derived essential feature analysis.
- Test results demonstrated that for highly imbalanced classes early fusion is the most promising method. Analysis of essential feature hadn't met the expectations.
- Stereo-mode classification results demonstrated significant improvement in comparison to neural network mono observations with background noise suppression close to the gamma/hadron imbalance observed in the experiment.

Thank you for your attention!